

Exams will be returned in lab, I will have scores posted as soon as I get them form your instructors.

Check out in lab will be held this week only.

If you fail to check out, you will be assessed a fine and have a hold placed on your records.

You last labs are due on the last day of the lab. Any labs turned in late will receive a ZERO score.

12/5/07 Dr. Mack. CSUS 2

Final Exam: Friday 12/21/07 8-10 am (here in lecture)

What will be covered on the exam?

Chapter 1 through 10

The format will be multiple choice.

Many of the questions will be the same as those on your exams!

You can redo the OWL exercises by going into the past due assignments for practice.

Isotope	Emission	Half-life	Applications
³ H	beta	12.3 years	To measure water content of body
³² ₁₅ P	beta	14.3 days	Detection of tumors, treatment of a form of leukemia
⁵¹ ₂₄ Cr	gamma	27.8 days	Diagnosis: size and shape of spleen, gastrointestinal disorders
⁵⁹ ₂₆ Fe	beta	45.1 days	Diagnosis: anemia, bone marrow function
60 27 Co	beta, gamma	5.3 years	Therapy: cancer treatment
⁶⁷ ₃₁ Ga	gamma	78.1 hours	Diagnosis: various tumors
⁷⁵ ₃₄ Se	beta	120.4 days	Diagnosis: pancreatic scan
⁸¹ ₃₆ Kr	gamma	2.1×10^5 years	Diagnosis: lung ventilation scan
85 38Sr	gamma	64 days	Diagnosis: bone scan
99 43 Tc	gamma	6 hours	Diagnosis: brain, liver, kidney, bone, and heart muscle scans
¹²³ ₅₃ I	gamma	13.3 hours	Diagnosis: thyroid cancer
¹³¹ ₅₃ I	beta, gamma	8.1 days	Diagnosis and treatment: thyroid cancer
¹⁹⁷ ₈₀ Hg	gamma	65 hours	Diagnosis: kidney scan
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Maintaining a critical distance helps to avoid over exposure.

Distance involves the use of the inverse square law of radiation:

The intensity of radiation is inversely proportional to the square of the distance from the source of the radiation.

$$I_{x} = I_{y} \times \frac{d_{y}^{2}}{d_{x}^{2}}$$

As the distance away increases (x) the intensity of radiation decreases. I_v and d_v are the measured intensities at a distance d_v .

12/5/07 Dr. Mack. CSUS 10

square of the distance from the source of the radia

hour? $I_{x} = I_{y} \times \frac{d_{y}^{2}}{d_{x}^{2}} \longrightarrow \frac{I_{x}}{I_{y}} = \frac{d_{y}^{2}}{d_{x}^{2}} \longrightarrow d_{x}^{2} = \frac{I_{y} \times d_{y}^{2}}{I_{y}}$

How far must one stay back to minimize exposure to 50 urads per

A radiation source is emitting at 100 mrads per hour at a distance of

3 ft.

12/5/07

$$d_{x} = \sqrt{\frac{100 \text{mrad} \times (3 \text{ft})^{2}}{50 \mu \text{rad} \times \frac{1 \text{mrad}}{10^{3} \mu \text{rad}}}} \qquad \qquad \bullet \qquad d_{x} = \sqrt{\frac{I_{y} \times d_{y}^{2}}{I_{x}}}$$

d_x =134 feet! Stay back Jack!

11

INDUCED NUCLEAR REACTIONS

Induced nuclear reactions are processes that take place when nuclei are bombarded with subatomic particles such as *alpha* particles or *neutrons*.

$${}^{14}_{7}N + {}^{1}_{0}n \longrightarrow {}^{14}_{6}C + {}^{1}_{1}p$$

An example of an induced nuclear reaction is the one that produces radioactive carbon-14 in the atmosphere.

This process takes place when a nitrogen-14 atom is struck by a cosmic ray neutron.

Induced nuclear reactions have been used to produce all of the elements in the periodic table with atomic numbers greater than 92.

Transuranium Elements

In 1940, the first of the transuranium elements—elements with a Z > 92—was synthesized by bombarding uranium-238 nuclei with neutrons

$${}^{238}_{92} U + {}^{0}_{1}n \rightarrow {}^{238}_{93} U$$

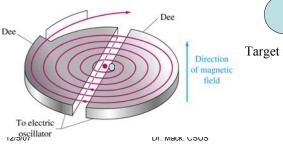
$${}^{239}_{92} U \rightarrow {}^{0}_{-1}e + {}^{239}_{93} Np$$

$${}^{239}_{93} Np \rightarrow {}^{0}_{-1}e + {}^{239}_{94} Pu$$

12/5/07 Dr. Mack. CSUS EOS 13

Transuranium Elements

Considerable energy must be imparted to a positive ion in order for it to overcome repulsion by a positively charged nucleus. A machine, called a *charged-particle accelerator*, or *cyclotron*, is capable of this process





Target nucleus

EOS

INDUCED NUCLEAR REACTIONS

Fission vs. Fusion

Heavy nuclei (Uranium) release energy when they split (Nuclear Fission)

The **product nuclei** mass is **less** than the sum of the parent nuclei.

Where does the mass go? Correct, ENERGY!

Light nuclei (hydrogen, helium) release energy when they **fuse (Nuclear Fusion)**

The **product nuclei** mass is **less** than the sum of the parent nuclei.

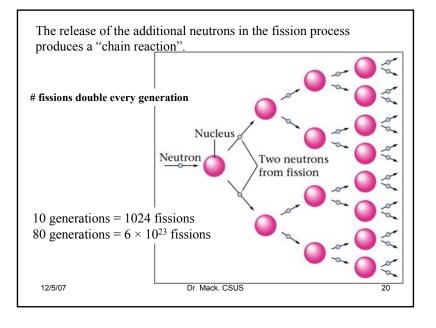
 $E = \Delta m \times c^2$

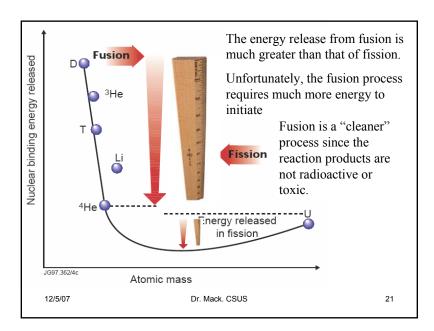
12/5/07 Dr. Mack. CSUS 16

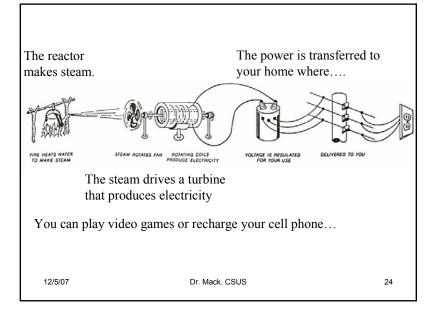
What is Nuclear Fusion?



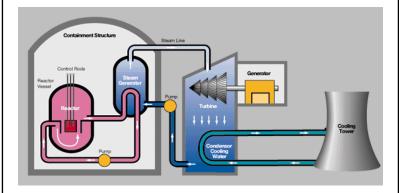
- Nuclear Fusion is the energy-producing process taking place in the core of the Sun and stars
- The core temperature of the Sun is about 15 million °C. At these temperatures hydrogen nuclei fuse to give Helium and Energy.







Pressurized Water Reactor



- The PWR has 3 separate cooling systems.
- Only 1 should have radioactivity, the Reactor Coolant System